



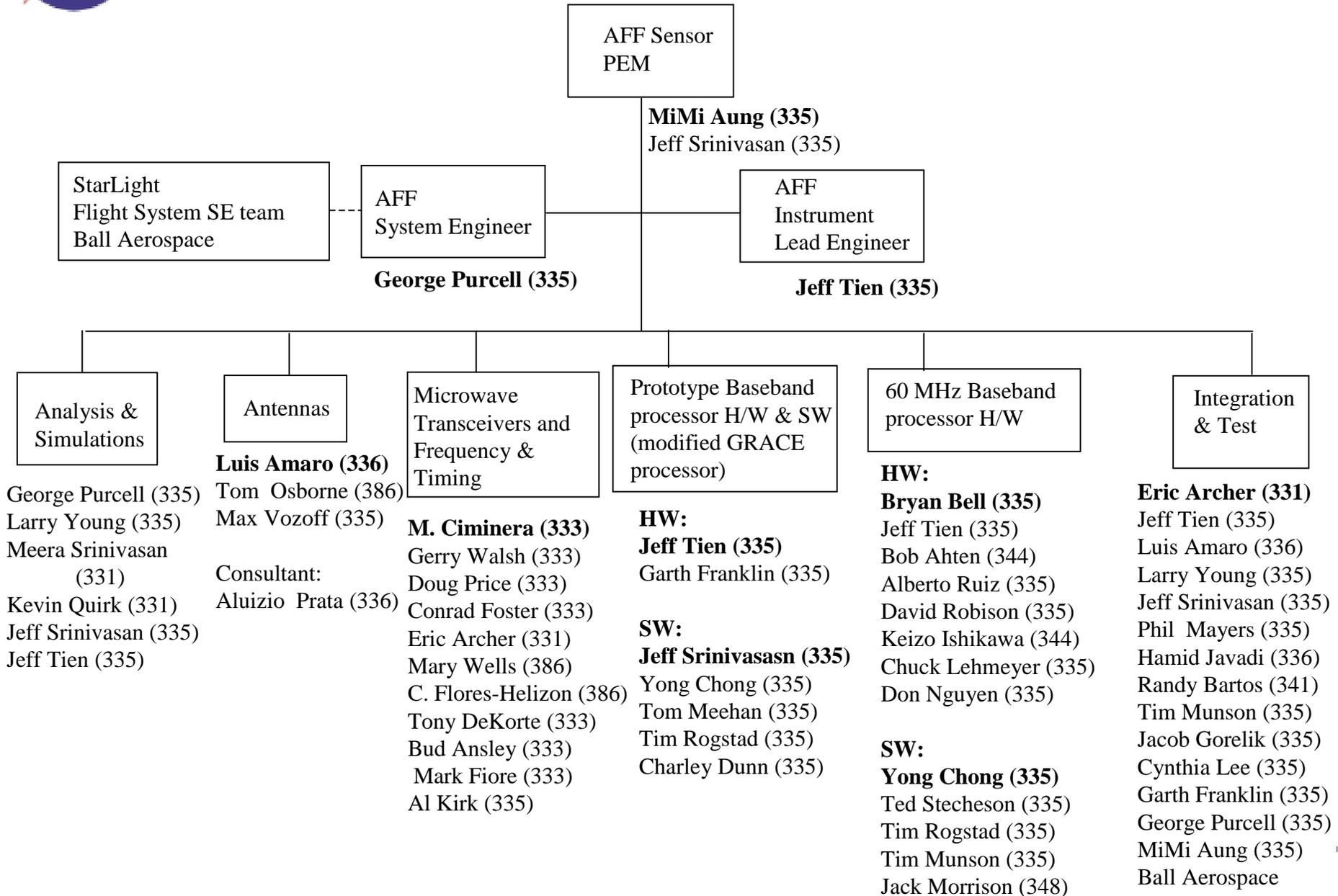
Autonomous Formation Flying (AFF) Sensor for Precision Formation Flying Missions



MiMi Aung
11/21/02



Contributors to the AFF Sensor





History and Status



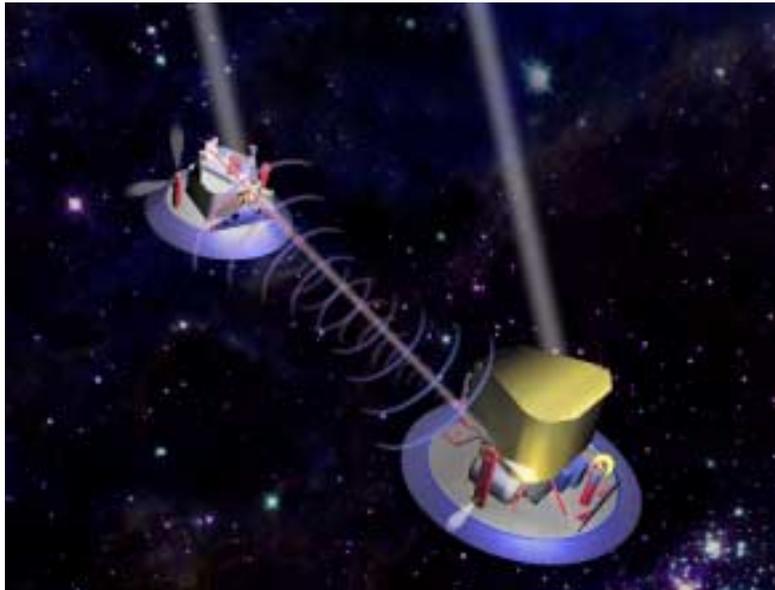
- AFF Sensor is a novel design innovated and patented by the JPL GPS team (335)
 - The AFF Sensor was initially "seeded" in a small exploratory technology task within the DSN Technology Program (now called the IND Technology Program in 9xx).
 - Infused through the New Millennium Program (NMP) for the DS-3 mission (Separated Spacecraft Interferometer).
 - Moved into the Origins Program where DS-3 -> ST-3 -> StarLight.
 - U.S. Patent No. 6,072,433, "An autonomous formation flying sensor for precise autonomous determination and control of the relative position and attitude for a formation of moving objects", June 6, 2000. (Lawrence E. Young, Stephen M. Lichten, Jeffrey Y. Tien, Charles E. Dunn, Bruce J. Haines, Kenneth H. Lau)
- Technology development activities 1999 - 2002 (StarLight project and Code R funding)
- At this time, a Ka-band prototype of the AFF Sensor has been developed and extensively characterized.
 - Fundamental algorithms have been demonstrated
- AFF Sensor is ready for adoption into future multiple spacecraft precision formation flying missions
 - With customization for individual missions.
- Being evaluated further under Terrestrial Planet Finder (TPF) pre-project technology program



AFF Sensor within a FF Mission



- The AFF Sensor is a radio-frequency sensor for multiple spacecraft precision formation flying (FF) missions. It provides:
 - Estimates of ranges and bearing angles among multiple spacecraft
 - A wide field of view for initial acquisition and lost-in-space scenarios.



StarLight: A separated spacecraft optical interferometer mission

StarLight key performance requirements	Directly facing (cone < 2°)	Nearly facing (2° < cone < 45°)	Not facing (cone > 45°)
*Range (cm)	2	2-30	160
*Bearing angles (arc-minute)	1	1-600	5400

*1- σ accuracy

Spacecraft separation:
Nominal: 30m – 1000m
Recovery capability: 1 - 10 km



Key Features



- Performance
 - (2 cm, 1 arcmin) accuracy when the spacecraft are directly facing each other
 - Wide field of view coverage ($\sim\pm 70^\circ$ cone)
 - 3-D relative positioning (range, azimuth angle, elevation angle)

- Autonomous
 - No real-time ground-based interaction
 - Self-contained instrument: Transmit, receive and data communication HW/SW on multiple spacecraft
 - No aid from Earth-based GPS system

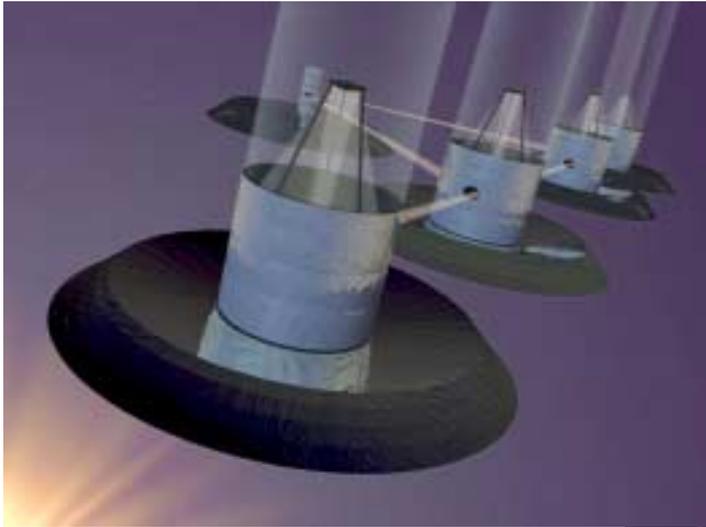
- Real-time
 - Real-time determination of range and bearing angles for real-time use in the formation flying control system



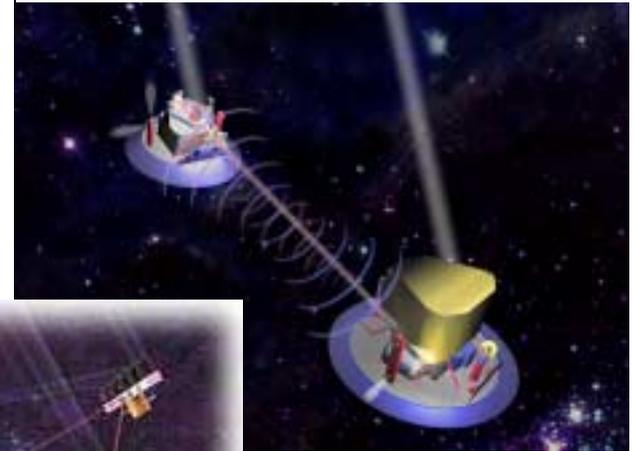
Examples of FF Missions



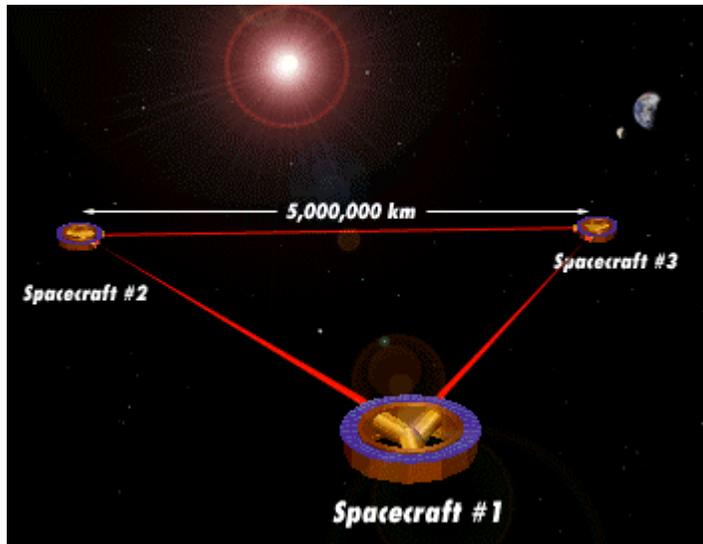
Terrestrial Planet Finder (TPF) ~ 2015



StarLight (flight portion cancelled)



Planet Imager (PI) ~20XX



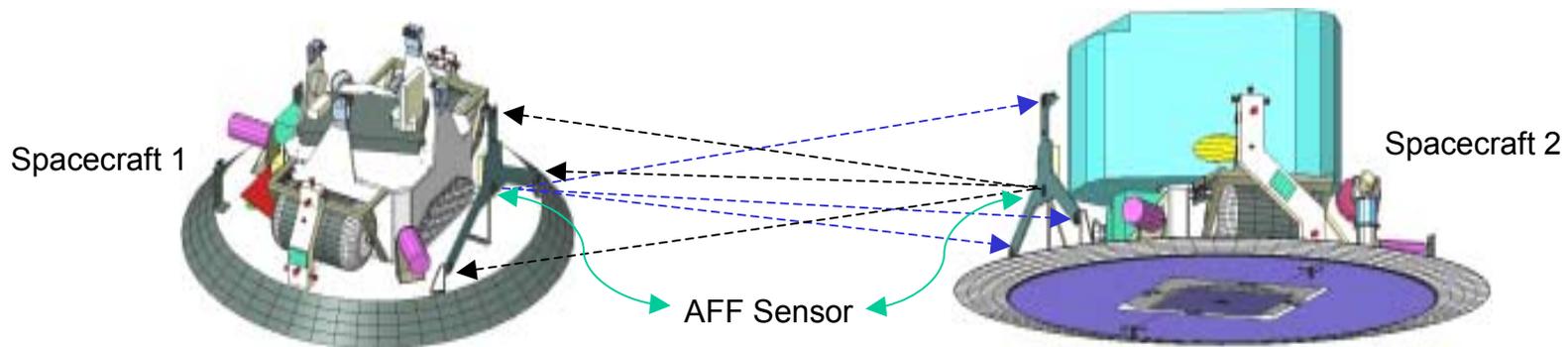
Laser Interferometer Space Antenna (LISA) ~ 2008



Design Description



- An RF instrument that is distributed over multiple s/c.
- AFF Sensor on each spacecraft transmits and receives GPS-like signals
$$S(t) = P(t)D(t)\cos(2\pi ft + \phi)$$
where $P(t)$ = ranging code
 $D(t)$ = Data bits (telemetry)
 f = carrier frequency (RF, Ka-band for StarLight)
- 1 TX and 3 RX on the front of each s/c (for determination of range and bearing angles)
- Range is derived mainly from ranging code delay between the s/c
- Bearing angles are derived mainly from carrier phase observables
- Telemetry exchanged on the RF link
 - Calibration across the two spacecraft
 - Enables each s/c to compute formation flying solutions

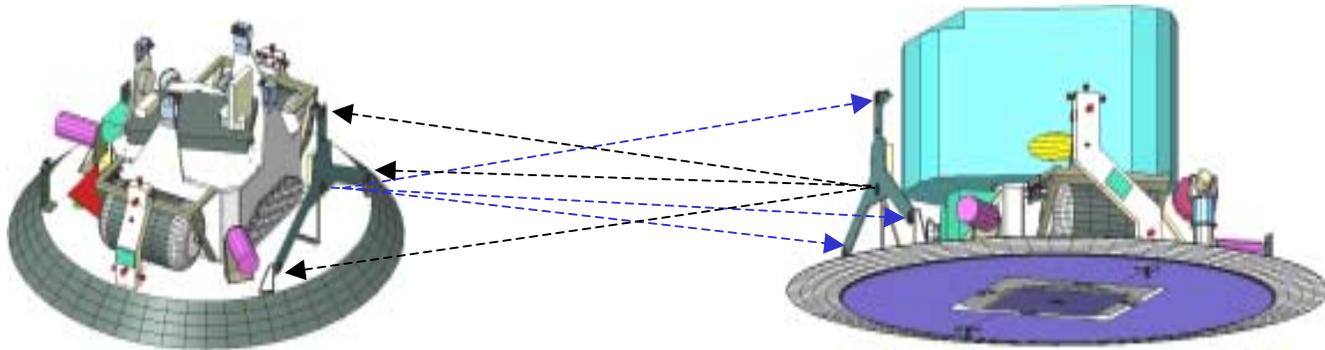




Design Description (Cont'd)

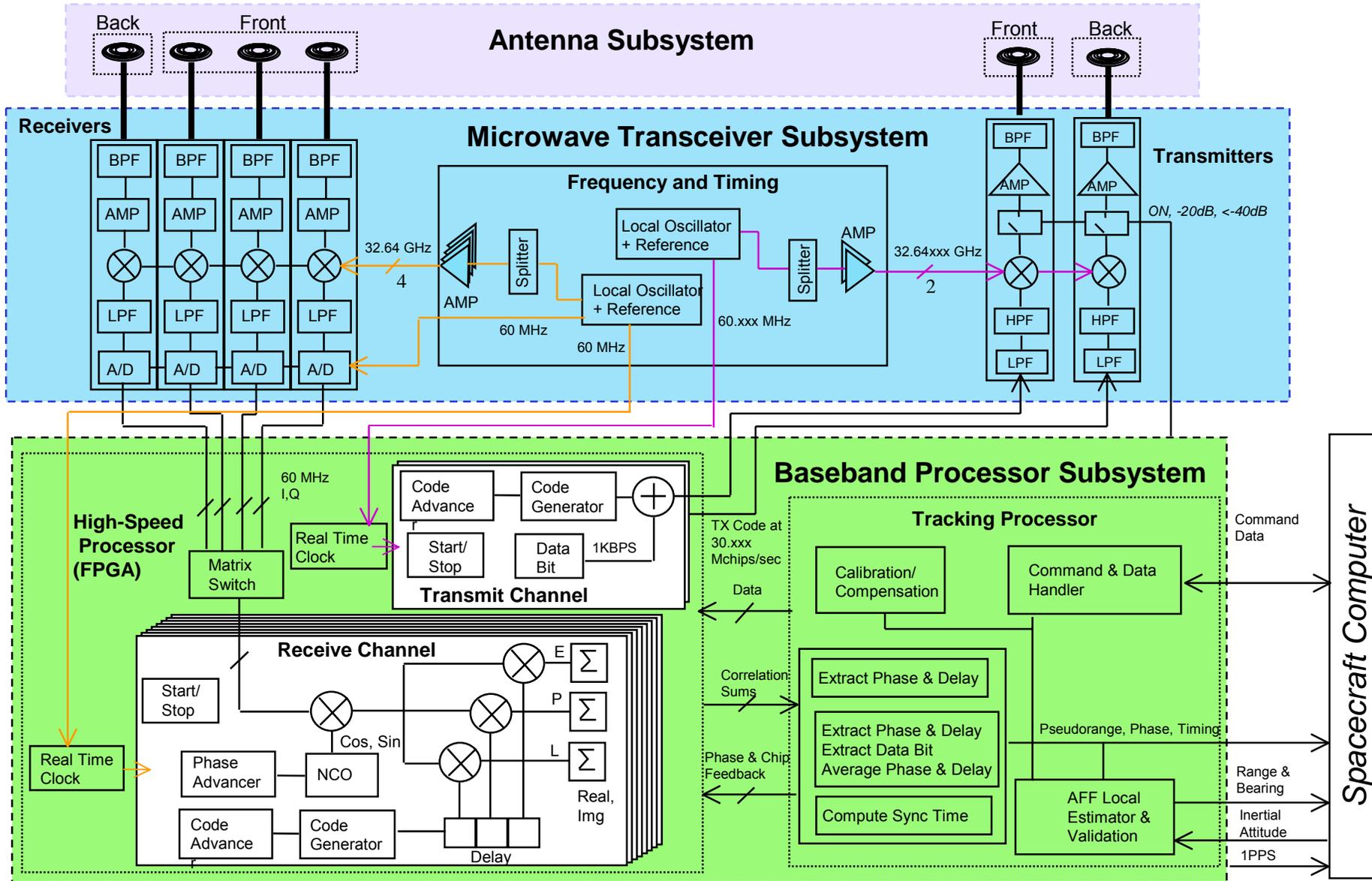


- Signal transmission and reception options
 - Simultaneously
 - Time-Division Duplexing (TDD)
 - Synchronously
 - Asynchronously





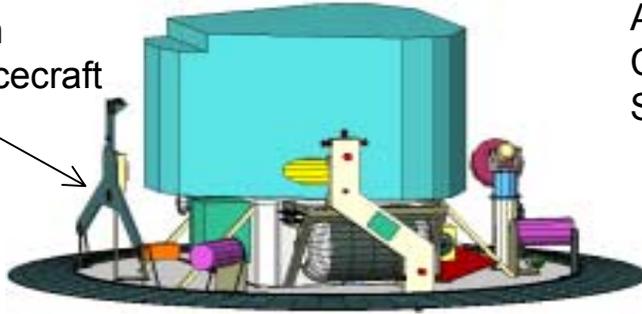
AFF Sensor Subsystems





Challenges in a Distributed S/C Mission **JPL**

AFF Sensor on
Combiner Spacecraft



AFF Sensor on
Collector
Spacecraft



Key challenges are:

- To achieve required RF performance in the presence of multipath
 - Effective antenna pattern
 - Effective isolation between TX and RX antennas
- To maintain insensitivity to thermal, electrical and mechanical instabilities
 - Continuous self-calibration techniques across multiple spacecraft
- To implement the required frequency scheme at Ka-band
- To operate as a single instrument distributed across multiple spacecraft
- Initial signal acquisition and calibration of the distributed system
- To be accommodated concurrently with other spacecraft subsystems and the interferometer, while minimizing multipath



Implementation Innovations



- Custom Antenna Design
 - To minimize multipath while keeping a wide field of view
- Ka-band implementation
 - Two closely spaced Ka-band references derived from a single (~10 MHz) reference source on each spacecraft
 - Coherence between RF signals with digital clocks
- Digital Signal Processing
 - Continuous, instantaneous, self-calibration scheme
 - Operates across distributed system
 - Removes clock offsets, instrumental variations
 - Carrier-aided smoothing algorithm to improve range estimates
 - Coherence of generated code with the RF signals

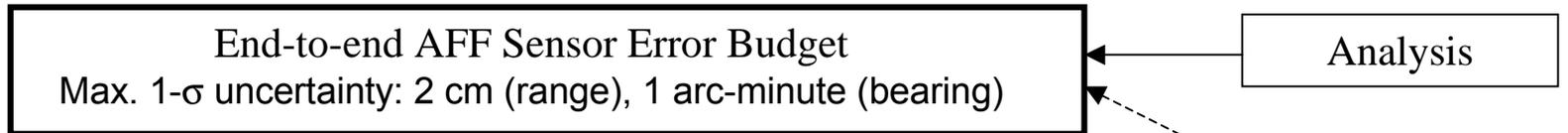


Technology Development



Key technology challenges have been addressed as follows:

- An end-to-end Ka-band prototype system was developed.
- Related spacecraft mockups were fabricated.
- Four testbeds were used.



Antenna pattern assessment Testbed



Outdoor Antenna Isolation Testbed



Indoor AFF Sensor Testbed



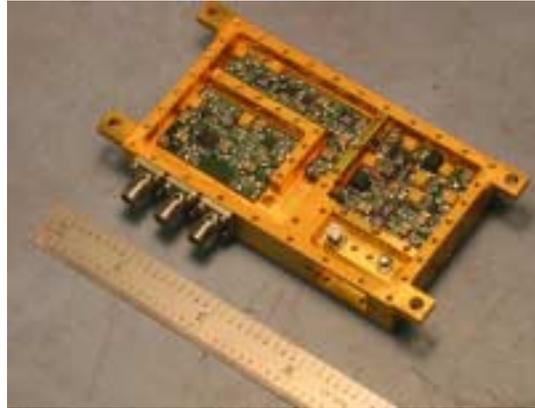
358-meter Range Outdoor Radiated Testbed



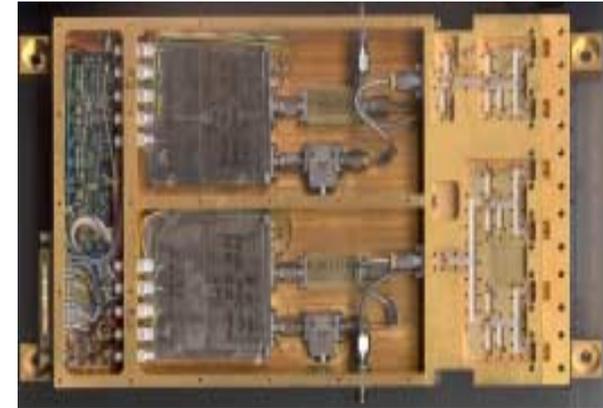
Prototype Hardware



Prototype Ka-band antenna with choke rings



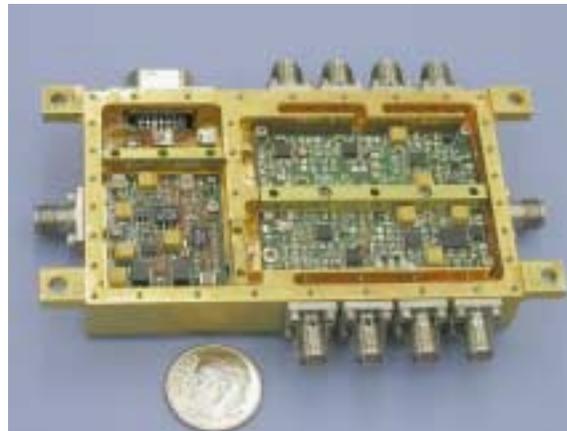
Ka-band Transmitter:
Output: 32.64 GHz RF signal at 13 dBm



Ka-band Local Oscillator:
Output: 32.64 GHz generated from 120 MHz input.



Prototype Baseband Processor – modified GRACE baseband processor (IPU)



Ka-band Receiver:
Input 32.64 GHz,
Output: 60 MHz 1-bit I and Q samples



Reference oscillator:
120 MHz

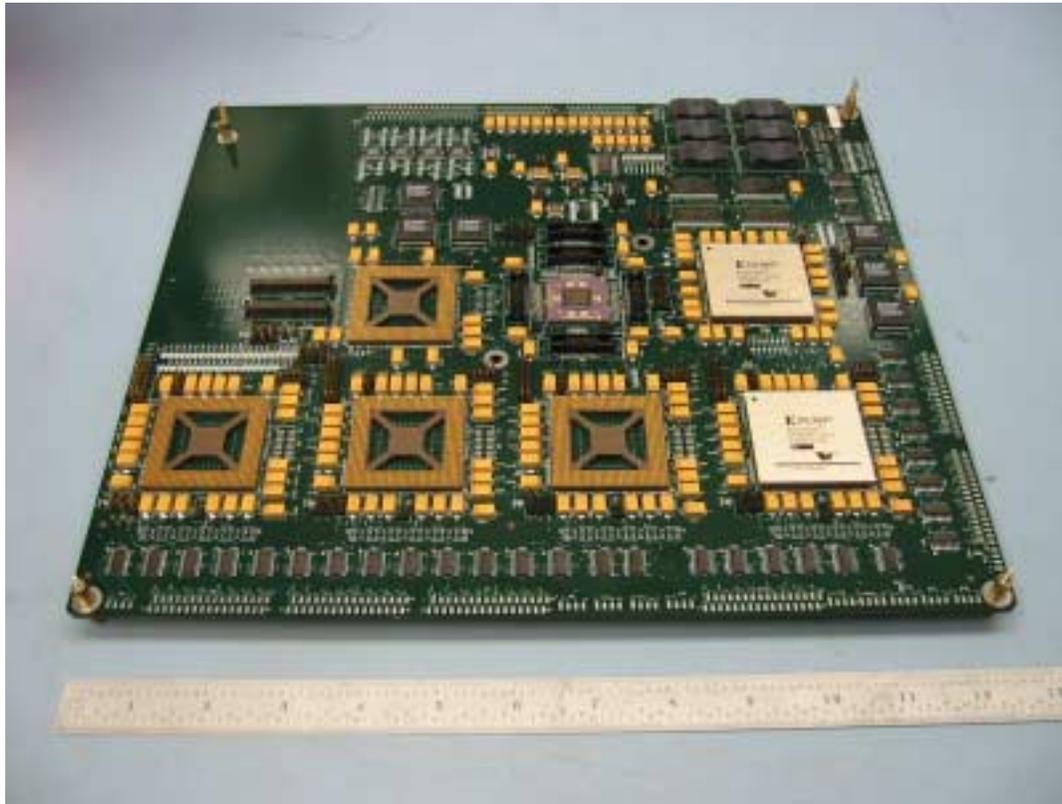


60 MHz Baseband Processor



A 60 MHz Baseband Processor will be completed in Q1 FY-03.

- Will provide more capability, flexibility and re-programmability for further investigation of the AFF Sensor.





AFF Sensor antennas mounted with mock-ups of the mounting plate and sunshade in the JPL 60-foot anechoic chamber

Objective:

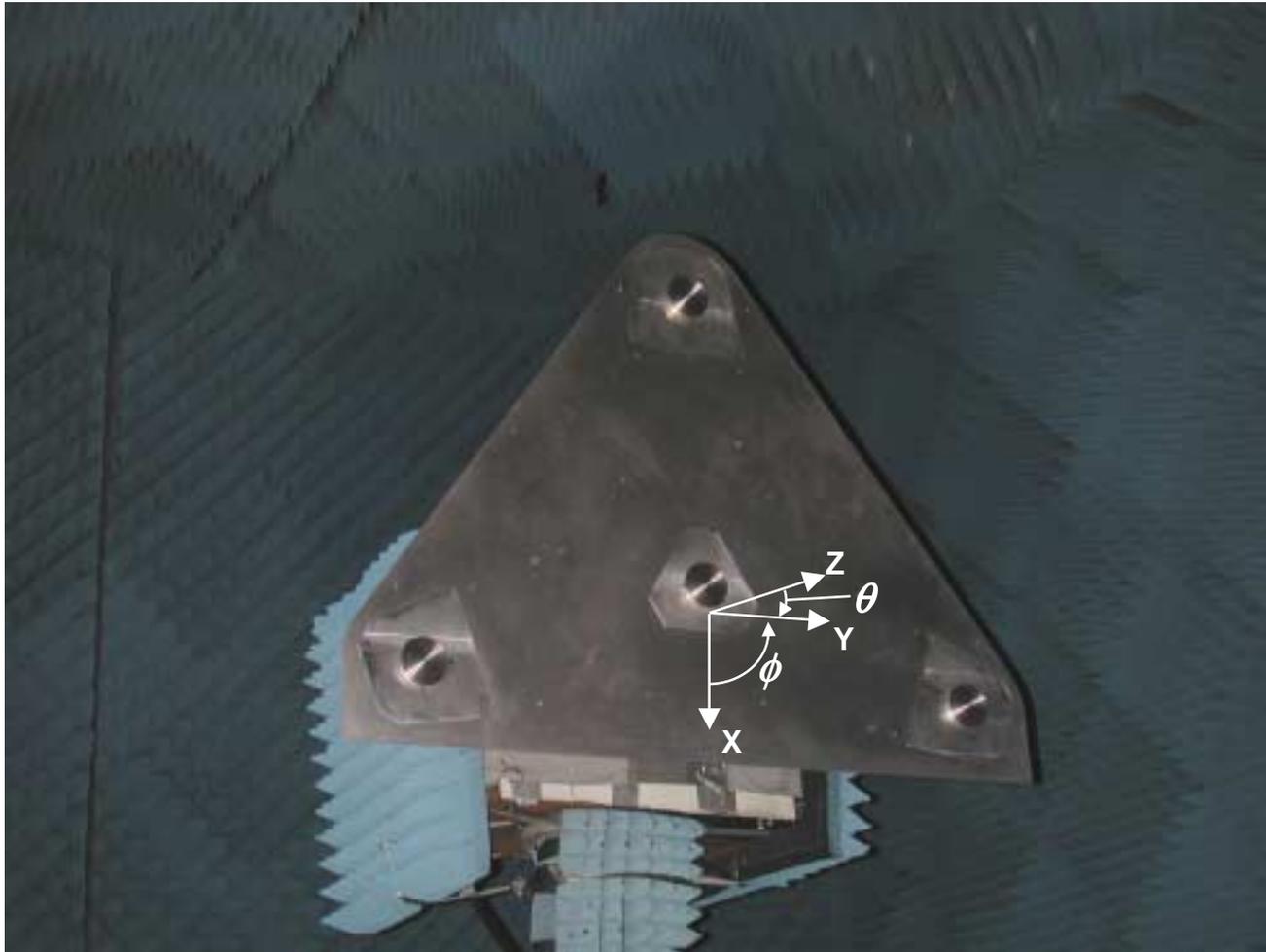
Evaluate degradation of the delay and phase patterns of the transmitting and receiving antennas due to spacecraft multipath sources.

Approach:

- Construct mockups of the AFF mounting plate and sunshades.
- Measure the gain and phase patterns of the antennas in the mocked-up flight environment.
- Compare measured antenna pattern deviations due to structural environment with the allocation within the end-to-end error budget.

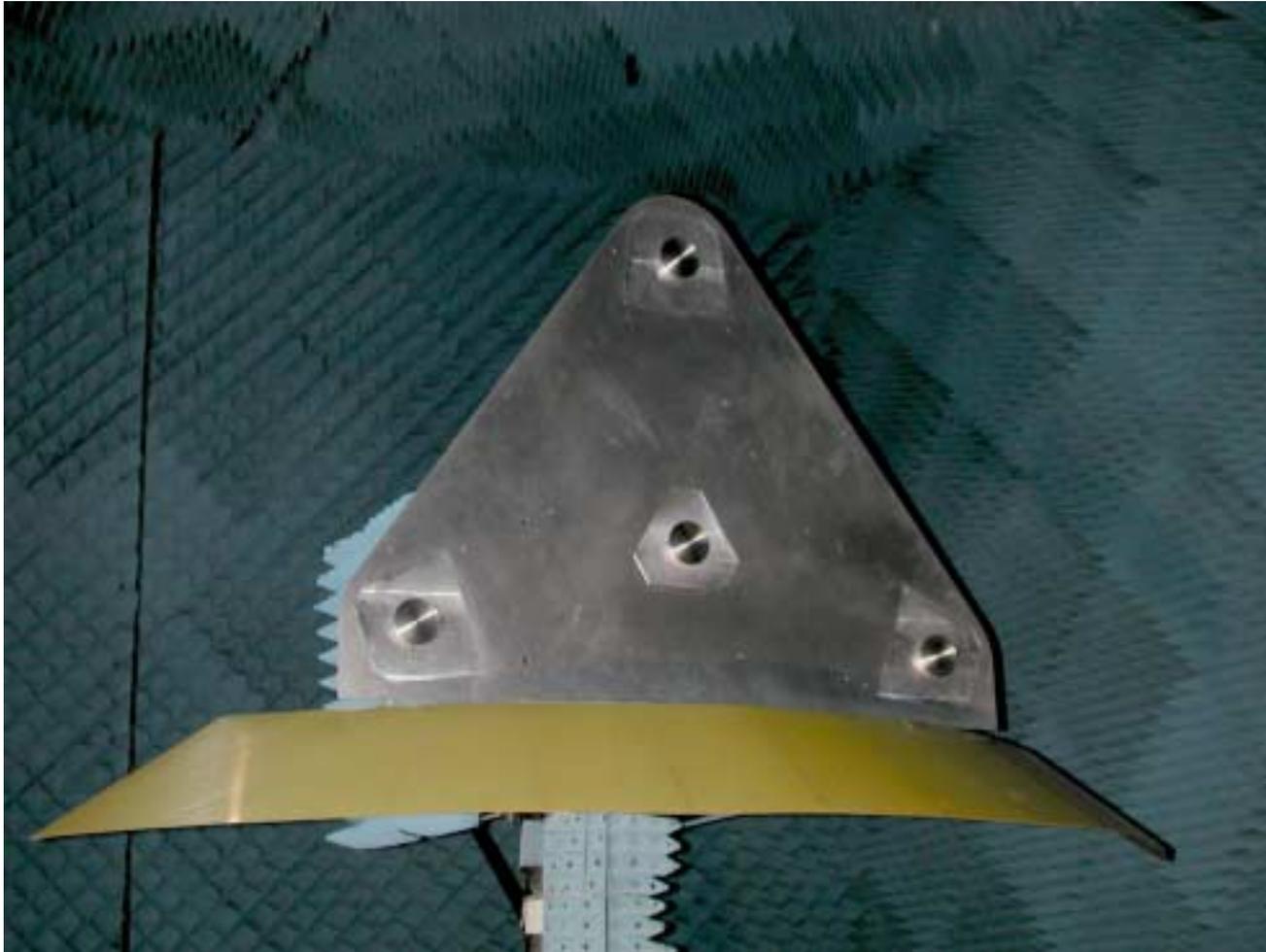


Antenna Pattern Testing (cont.) - Antenna Plate Baseline



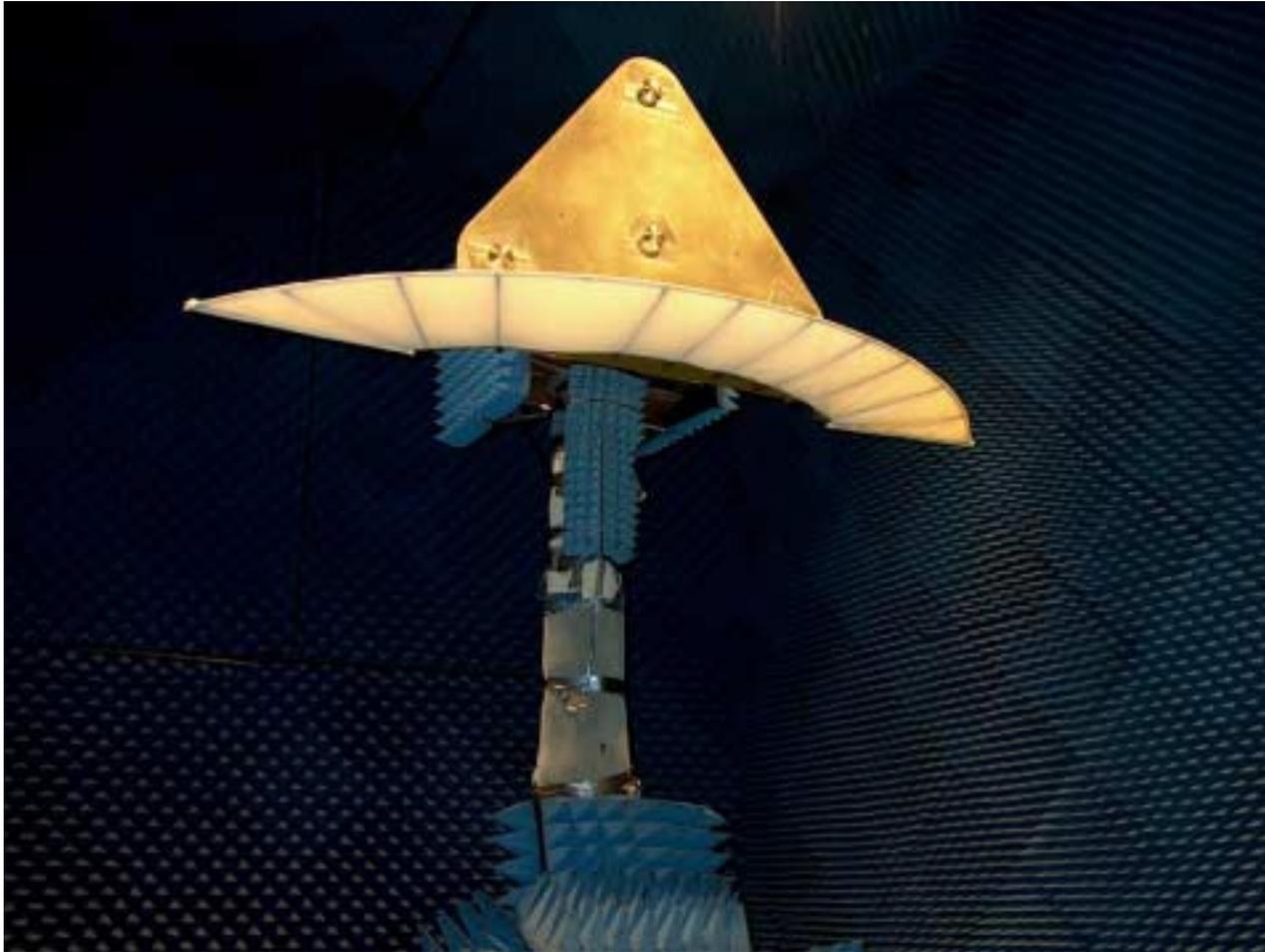


Antenna Pattern Testing (cont.) – Collector Shade



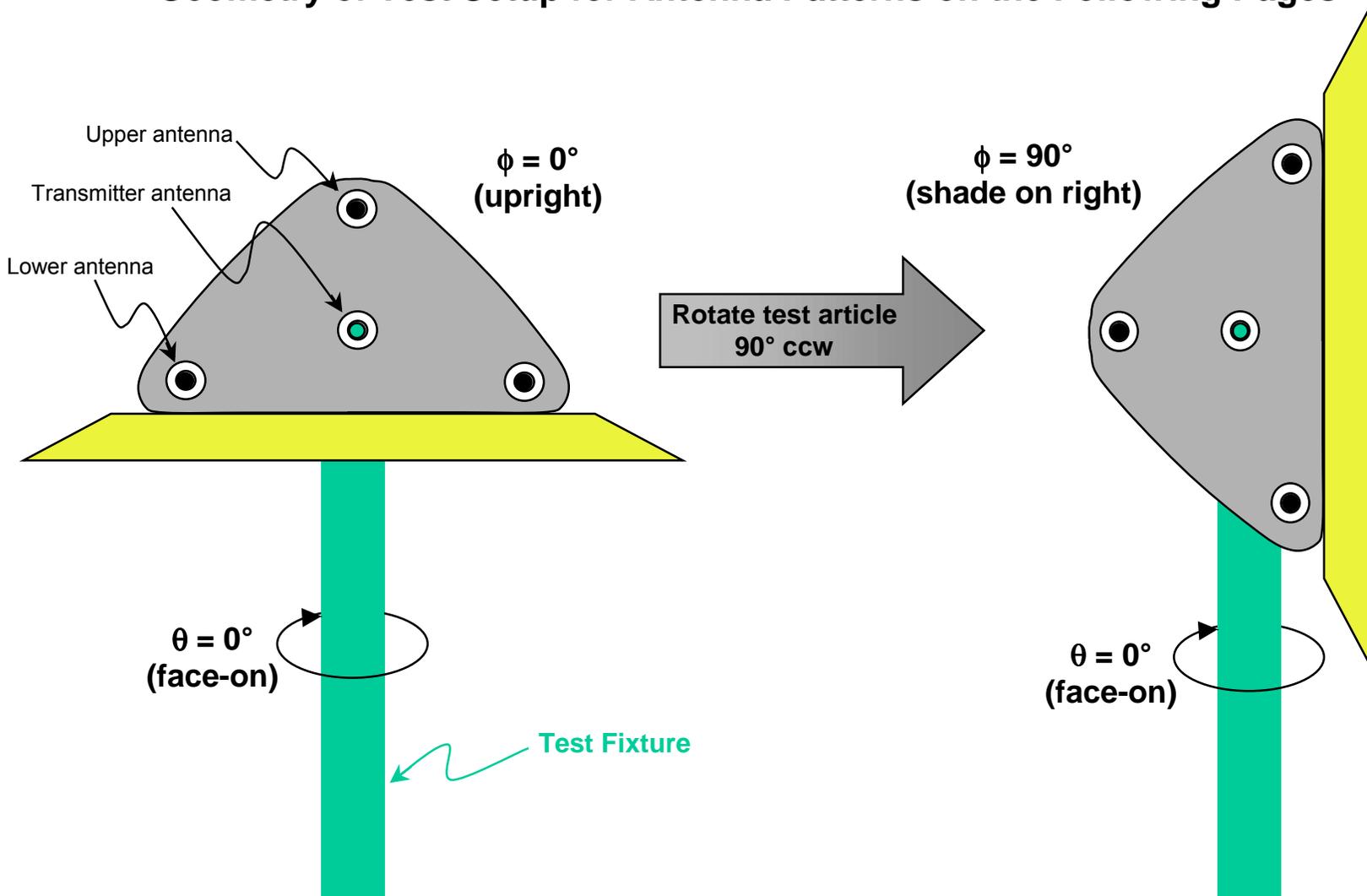


Antenna Pattern Testing (cont.) – Combiner Shade

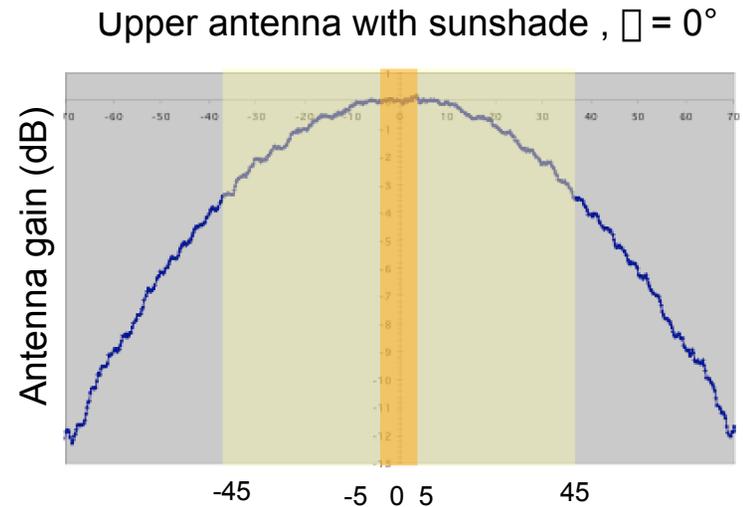
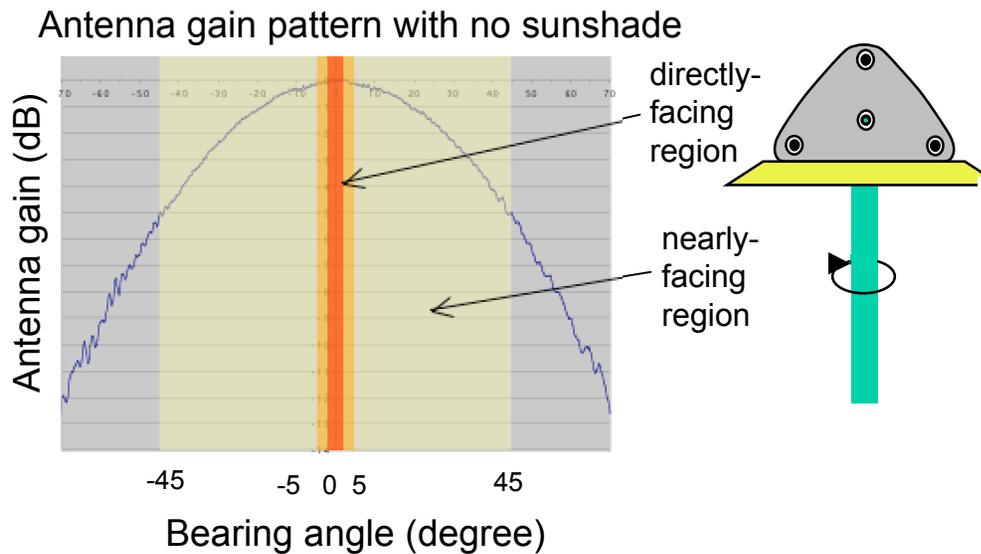




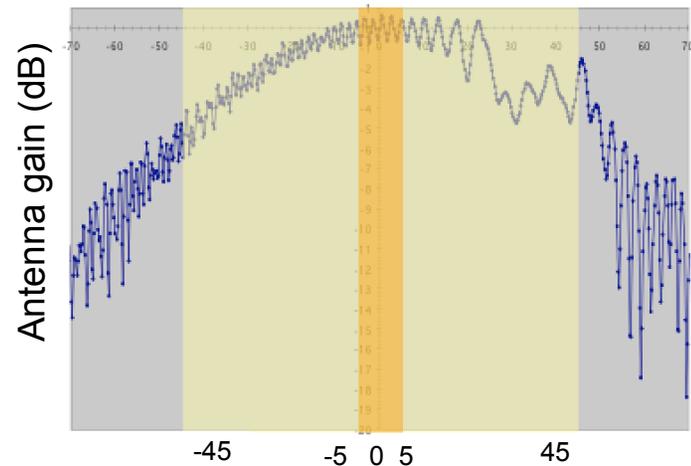
Geometry of Test Setup for Antenna Patterns on the Following Pages



AFF Sensor Antenna Pattern Assessment Testbed (Cont'd)

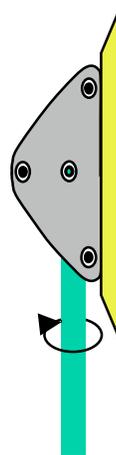


Lower Antenna gain pattern with sunshade, $\alpha = 90^\circ$



Conclusion

- Antenna pattern is degraded by the sunshade.
- Deviations from the nominal pattern (uncalibrated errors) fed into error trees show that the AFF Sensor can still meet the (2 cm, 1 arc-min) requirement
 - Degrades slower than requirement relaxation away from boresight.





Objective:

Determine whether isolation between the transmitting and receiving antennas on the same spacecraft is sufficient and stable.

Approach:

- Construct mockups of the mounting plate and sun shades.
- Measure isolation between the antennas with and without mocked-up flight environment.



With no sunshade



With Collector s/c sunshade



With Combiner s/c sunshade



Conclusion:

- Without the sunshade, the measured levels of isolation matched predicted levels.
 - Antenna mounting plate did not introduce any unpredictable effects.
- Sunshade degraded isolation levels.
 - Level of degradation varied with the shape of the sunshade and with changes in location of the sunshade.
 - Repeatability is poor due to effects at the small Ka-band wavelengths.
 - Multipath sources are localized.
- Possible to control isolation levels by placement of absorber at strategic locations.
- Consider Time-division duplexing (TDD) scheme on individual mission basis.





Objective:

Verify fundamental algorithms distributed across multiple spacecraft.

Approach:

- Integrate an indoor testbed representative of the AFF Sensor distributed on two “spacecraft.”
- Composed of Ka-band and digital modules on two sides connected by adjustable waveguide attenuators representative of the space loss.
- Each half of the sensor is operated from an independent frequency reference.

AFF Sensor on
Spacecraft #1

representative space-loss

AFF Sensor on
Spacecraft # 2

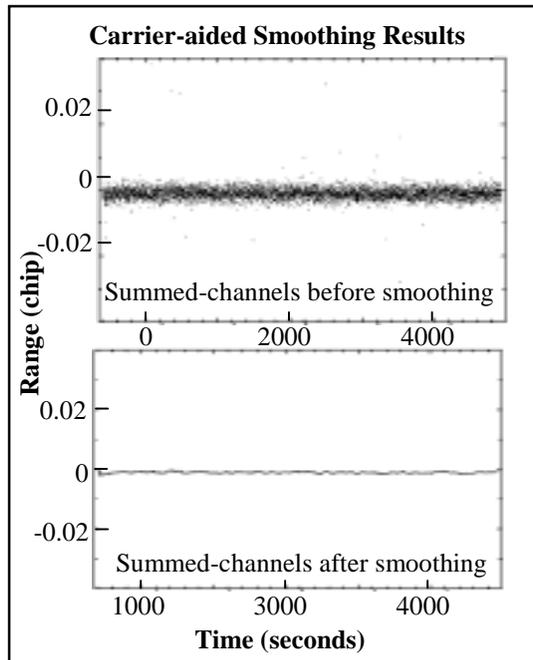




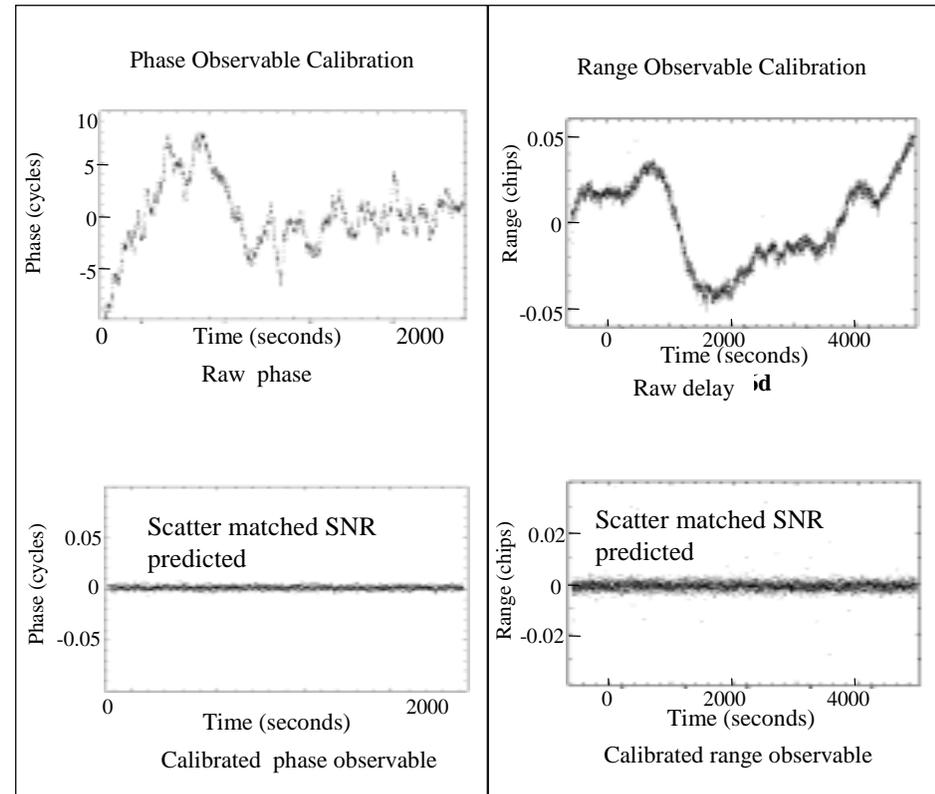
AFF Sensor Indoor Testbed (Cont'd)



- Fundamental, distributed Sensor algorithms have been verified.
 - Distributed operation
 - Ka-band scheme
- Carrier-aided smoothing algorithm was verified.



- Continuous self-calibration across two halves was verified.
 - Phase observable
 - Range observable





AFF Sensor Indoor Testbed



Conclusion:

- The following key technologies were demonstrated in the distributed environment:
 - Fundamental AFF Sensor scheme
 - Continuous self-calibration algorithm operating across two independent halves
 - Carrier-aided smoothing algorithm requiring sustained coherence across each spacecraft
 - Basic Ka-band scheme supporting the Sensor design
 - Time-Division Duplexing (TDD) scheme



End-to-End Functionality Field Test across a 1200-foot Outdoor Range



Objective:

- Verify end-to-end functionality of the complete AFF Sensor. (Full performance is not expected in the presence of uncontrolled multipath sources in the outdoor environment.)

Approach:

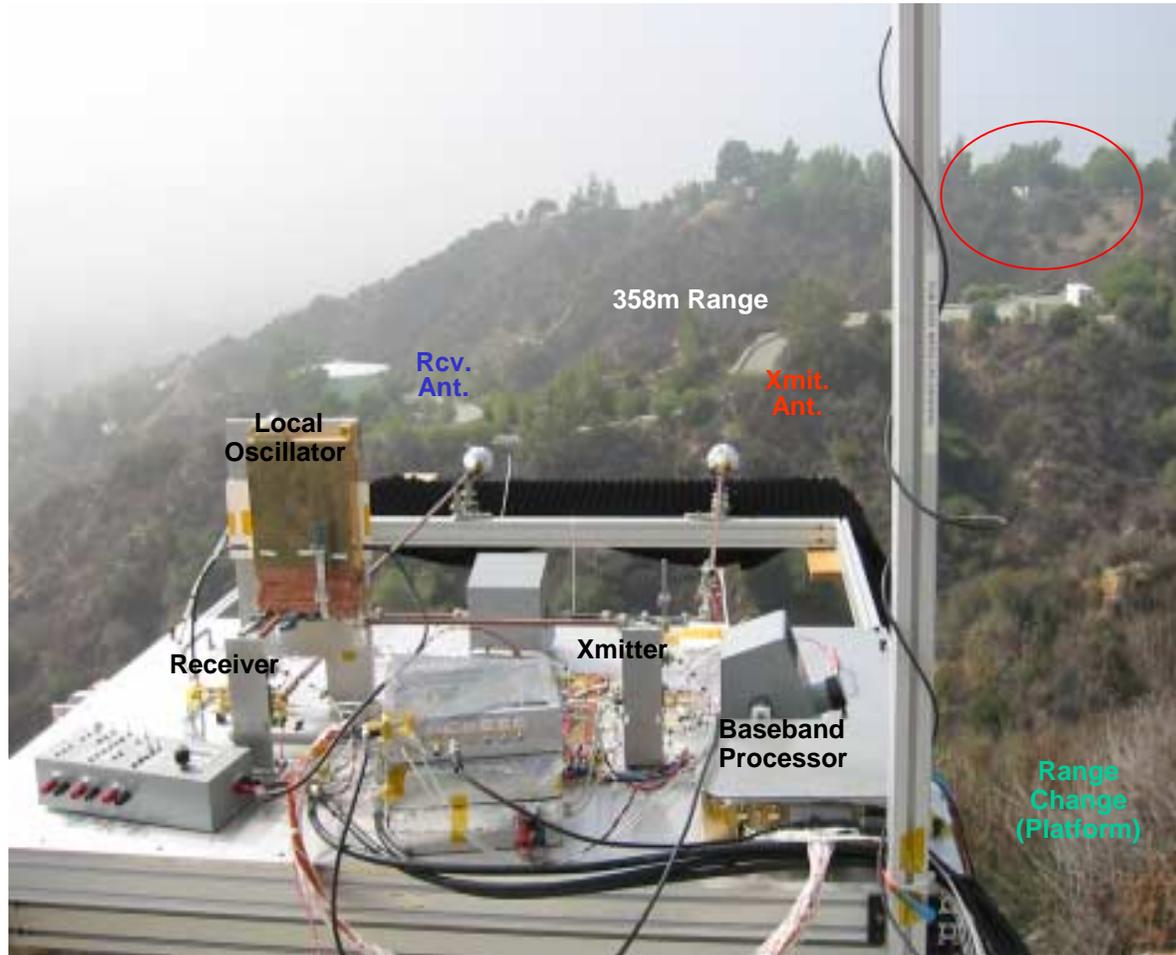
- Operate the prototype AFF Sensor distributed over two halves across a 1200-foot outdoor range.
- Introduce changes in ranges and bearing angles.
- Derive estimates of the range and bearing angle from observables measured during end-to-end operation.



End-to-End Functionality Field Test across a 1200' Outdoor Range (Cont'd)



East End





End-to-End Functionality Field Test across a 358-m Outdoor Range (Cont'd)



West End



Autonomous Formation Flying (AFF) Sensor



End-to-End Functionality Field Test across a 358-m Outdoor Range (Cont'd)



Conclusion:

- End-to-end functionality of the AFF Sensor has been verified by successful operation across the 1200-foot range.
 - Range, range-change and bearing angle were determined successfully.
 - Measured ranges matched the GPS-surveyed “truth” ranges (within the accuracy expected in the presence of uncontrolled multipath)
 - Range-change and bearing angle estimates matched the “truths” in the experiment.
 - Full end-to-end performance needs to be determined by operation across a large (>30 m) range with space-like conditions.



Conclusion



- A prototype of the AFF Sensor is fully functional.
- Fundamental algorithms have been verified for operation in a distributed spacecraft environment.
- Performance dependence upon the spacecraft architecture is understood.
- Results show that AFF Sensor can meet the StarLight requirements.
- Is ready for adoption into future multiple spacecraft precision formation flying mission
 - Sensor providing coverage from lost-in-space to full performance at face-to-face spacecraft configuration
 - Real-time
 - Autonomous
 - Applicable in deep space, near-Earth or regions with no access to GPS
 - Flexible FPGA-based signal processing
 - Can be augmented with star-trackers, Global Positioning System receivers (for near-Earth application)
 - For each mission, optimize on individual basis by design trade-off among: spacecraft design, Sensor field of view, formation flying system, instrument design.



Conclusion (Cont'd)



Further Investigations for Application to TPF

- Extend for five-spacecraft sensor design
 - Simultaneous multiple links in a dynamic environment
 - Which spacecraft are sensing signal from which other spacecraft under what circumstances
 - Antenna configuration
- Requires instantaneous 4π steradian coverage
- Much tighter non-directly facing requirements
 - Multipath modeling and mitigation
 - Self-Jamming (evaluate TDD for five S/C)
 - Near-Far issue (jamming from other S/C)
- New signal structure to avoid spacecraft rotation maneuver to resolve bearing angle ambiguities
- Integrated inter-spacecraft communications



Back-up



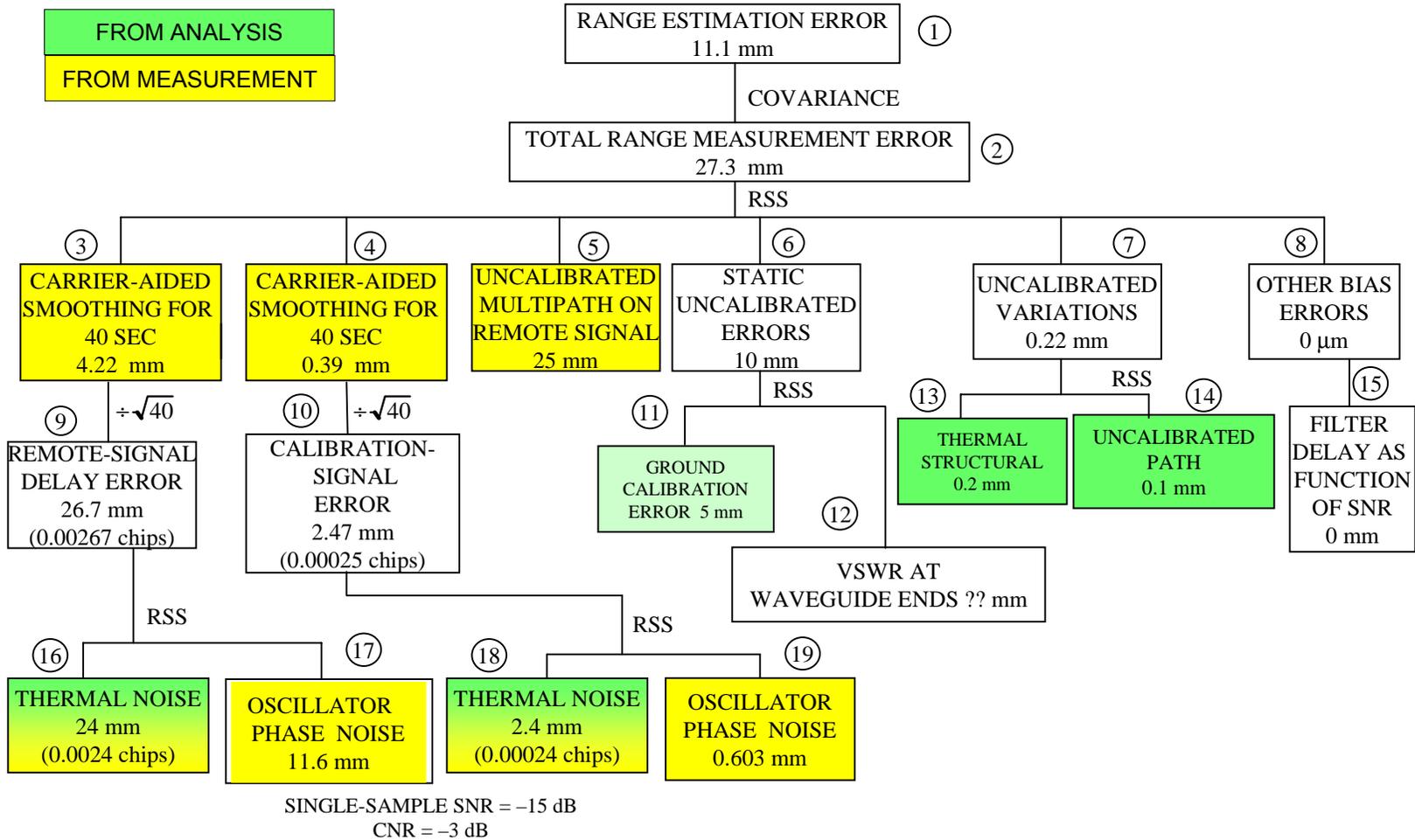
Key Parameters



PARAMETER	VALUE
RF carrier frequency	32.64 GHz
Chip rate of the PRN ranging code	30 Mchips/s
Sample rate	60 Msamples/s
Transmitted power	20 mW (13 dBm)
Transmitting and receiving antenna gain, on axis	9.2 dBi
Polarization loss (transmitting linear, receiving circular)	3 dB
Sky background temperature	3 K
Receiver noise temperature	2030 K
Receiver noise bandwidth	80 MHz
Separation between spacecraft	30-1000 m



TIME DIVISION DUPLEXING RANGE ERROR BUDGET, ANY CALIBRATION STATE, range 1000 m, 1-second observables





TIME DIVISION DUPLEXING BEARING-ANGLE ERROR BUDGET, GROUND CALIBRATION + IN-ORBIT ROTATION CALIBRATION, range 1000 m, 1-second observables

